

Inclined magnetic field effect on natural convective slip flow through a porous channel with convective boundary

K. Madhusudhan Reddy^{1*}, K. Kaladhar², D. Srinivasacharya²

¹Department of Mathematics, NIT Puducherry, Karaikal, India

²Department of Mathematics, NIT Warangal, India

Corresponding Author: K. Madhusudhan Reddy (madhu.nitpy@gmail.com)

ABSTRACT

In this paper, we consider the inclined magnetic field, Hall current and convective boundary on laminar incompressible natural convection flow through a porous channel with slip conditions. The governing equations determined the flow regime is transformed into a system of ordinary differential equations by utilizing suitable similarity transformations. Spectral Quasilinearization Method (SQLM) has been applying to solve the dimensionless governing equations; those were received by using similarity transformations from the system of governing partial differential equations with boundary conditions. The influence of emerging parameter on fluid flow velocities and temperature are presented graphically.

KEYWORDS: - Inclined magnetic field, Hall current, Convective boundary, Porous medium, SQLM

I. INTRODUCTION

Natural convection flow with heat transfer in a porous channel has an incredible significance in various fields. Natural convection has attracted a great deal of attention from researchers because of its presence both in nature and engineering applications. In nature, convection cells formed from air rising above sunlight-warmed land or water are a major feature of all-weather systems. The significance and developments of heat transfer have been addressed by many researchers [1, 2]. Hatami and Ganji [3] studied the natural convection of sodium alginate (SA) non-Newtonian nanofluid flow between two vertical flat plates by analytical and numerical methods. Recently, Sheremet et.al [4] presented the natural convection in an inclined cavity with time-periodic temperature boundary conditions using nanofluids: application in solar collectors. Most recently, Dogonchi et.al [5] studied the nature of magnetohydrodynamic natural convection of Cu-water nanofluid in a wavy cavity using CVFEM numerically.

An inclined magnetic field is a magnetic field with a non-zero inclination, has gained the attention of many researchers [6, 7]. Selimefendigil and Oztop [8] analyzed the natural convection in a flexible sided triangular cavity with internal heat generation under the effect of an inclined magnetic field. Al-Rashed et.al [9] presented the effects of external magnetic field inclination on laminar natural convection heat transfer in CNT-water nanofluid filled cavity. Recently, Atashafrooz et.al [10] discussed the interaction effects of an inclined magnetic field and nanofluid on forced convection heat transfer and flow irreversibility in a duct with an abrupt contraction.

Makinde and Aziz [11] studied the boundary layer flow of a nanofluid past a stretching sheet with a convective boundary condition. Oyelakin et.al [12] presented the unsteady Casson nanofluid flow over a stretching sheet with thermal radiation, convective, and slip boundary conditions. Khan et.al [13] analyzed heat transfer effects on carbon nanotube suspended nanofluid flow in a channel with non-parallel walls under the effect of velocity slip boundary condition: a numerical study. Recently, Ellahi et.al [14] investigated the effects of MHD and slip on heat transfer boundary layer flow over a moving plate based on specific entropy generation. Most recently, Chemetov et.al [15] presented the weak-strong uniqueness for fluid-rigid body interaction problems with slip boundary conditions.

In this article, the free convection Navier slip flow in a porous channel under an inclined magnetic field, Hall effects are investigated. A Spectral quasilinearization method is employed to solve the system of equations. This method was suggested by Bellman et.al [16] as an implication of the Newton-Raphson method. Ahmad and his co-workers [17, 18] have been used the quasilinearization method for Neumann and mixed boundary value problems. [19, 20] have been elongated the significance of the quasilinearization method to a broad variety of nonlinear BVP's and fractional differential equations and its applications. The exactness and capability of the Spectral quasilinearization scheme are explained by [21].

II. FORMULATION OF THE PROBLEM

Consider a steady incompressible, laminar free convection flow along the porous channel with distance $2d$ apart. A uniform external magnetic field B_0 is applied in the direction which makes an angle α with the positive direction of the x -axis, temperatures and concentrations are considered as shown in Figure 1. As the boundaries are infinitely extended in the x -direction, without loss of generality, we considered that the physical parameters are functions of y only. The properties of the fluid are presumed to be constant except for density variations in the buoyancy force term. With the above assumptions and Boussinesq approximations, the governing equations for the flow are given by

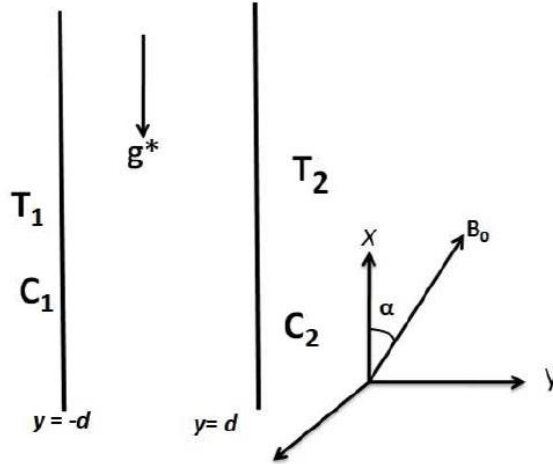


Fig. 1: Classical representation of the coordinate system.

$$\frac{\partial v}{\partial y} = 0 \Rightarrow v = v_0 \quad (1)$$

$$\rho v_0 \frac{\partial u}{\partial y} = \mu \frac{\partial^2 u}{\partial y^2} + \rho g^* [\beta_T (T - T_1)] - \frac{\sigma B_0^2 \cos \alpha}{1+m^2 \cos^2 \alpha} [u \cos \alpha - v_0 \sin \alpha + m w \cos^2 \alpha] - \frac{\epsilon \mu}{K_f} u \quad (2)$$

$$\rho v_0 \frac{\partial w}{\partial y} = \mu \frac{\partial^2 w}{\partial y^2} + \frac{\sigma B_0^2 \cos \alpha}{1+m^2 \cos^2 \alpha} [m u \cos \alpha - m v_0 \sin \alpha - w] - \frac{\epsilon \mu}{K_f} w \quad (3)$$

$$\rho C_p v_0 \frac{\partial T}{\partial y} = \frac{\sigma B_0^2}{1+m^2 \cos^2 \alpha} [(u \cos \alpha - v_0 \sin \alpha)^2 + w^2 \cos^2 \alpha] + K_f \frac{\partial^2 T}{\partial y^2} + \mu \left[\left(\frac{\partial u}{\partial y} \right)^2 + \left(\frac{\partial w}{\partial y} \right)^2 \right] \quad (4)$$

With

$$y = -d: u = w = 0, k \frac{\partial T}{\partial y} = -h_1 (T - T_1); y = d: u = w = 0, k \frac{\partial T}{\partial y} = -h_2 (T - T_2) \quad (5)$$

where u , v and w are the velocities in x , y and z respectively, μ is the coefficient of viscosity, g^* is the acceleration due to gravity, K_f is the thermal diffusion ratio, ρ is the density, C_p is the specific heat, $m = \eta_1 \sigma B_0$ is the Hall parameter, η_1 is the Hall factor, β_T is the coefficient of thermal expansion, h_1 and h_2 are the heat and mass transfer coefficients.

Introducing the following transformations

$$y = \eta d, u = \frac{\gamma Gr}{d} f, w = \frac{\gamma Gr}{d} g, T - T_1 = (T_2 - T_1) \theta \quad (6)$$

Substitute in Eqs. (2) - (4), we obtain the governing dimensionless equations as

$$f'' - Re f' + \theta - \frac{Ha^2 \cos \alpha}{1+m^2 \cos^2 \alpha} [f \cos \alpha - \lambda \sin \alpha + m g \cos^2 \alpha] - \frac{\epsilon}{Da} f = 0 \quad (7)$$

$$g'' - Re g' + \frac{Ha^2 \cos^2 \alpha}{1+m^2 \cos^2 \alpha} [m f \cos \alpha - g - m \lambda \sin \alpha] - \frac{\epsilon}{Da} g = 0 \quad (8)$$

$$\theta'' - Re Pr \theta' + \frac{Br Ha^2}{1+m^2 \cos^2 \alpha} [(f \cos \alpha - \lambda \sin \alpha)^2 + g^2 \cos^2 \alpha] + Br Gr^2 [(f')^2 + (g')^2] = 0 \quad (9)$$

With

$$f(-1) = g(-1) = \theta'(-1) - Bi_1(-1) = 0, \quad f(1) = g(1) = \theta'(1) - Bi_2(1) = -Bi_2 \quad (10)$$

where the primes indicate the differentiation concerning η , $Re = \frac{\rho v_0 d}{\mu}$ is Reynolds number, $Gr = \frac{g^* \beta_T (T_2 - T_1) d^3}{\nu^2}$ is thermal Grashof number, $Pr = \frac{\mu C_p}{K_f}$ is Prandtl number, $Br = \frac{\mu v^2}{K_f d^2 (T_2 - T_1)}$ is Brinkman number, $Ha = dB_0 \sqrt{\frac{\sigma}{\nu}}$ is the magnetic parameter, $Bi_1 = \frac{h_1 d}{k}$, $Bi_2 = \frac{h_2 d}{k}$ are the Biot numbers, $\lambda = \frac{Re}{Gr}$, $Da = \frac{K_f}{d^2}$ is the Darcy number, and m is the Hall parameter.

III. DISCUSSION OF RESULTS

The flow Eqs. (7) - (9) subject to the boundary conditions(10) are nonlinear and coupled, hence the system of equations is solved numerically using the Spectral quasilinearization method as explained in the works of [24]. The influence of the Biot number (Bi) on $f(\eta)$, $g(\eta)$ and $\theta(\eta)$ can be shown in figures 2-4 at $Pr=0.71, Re=2, Gr=2, \epsilon=0.1, \alpha=\pi/3, Br=3, Da=0.2, Ha=3, \beta_1=0.01$ and $\beta_2=0.01$. It is noted from figures 2-4 that the velocities and temperature profile decreases with an increase in Biot number (Bi).

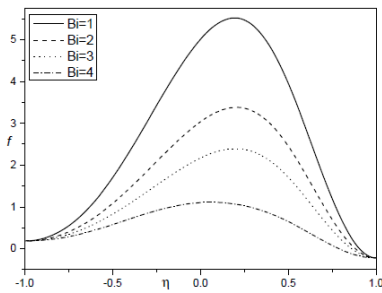


Fig 2. Effect of Bi on x -velocity

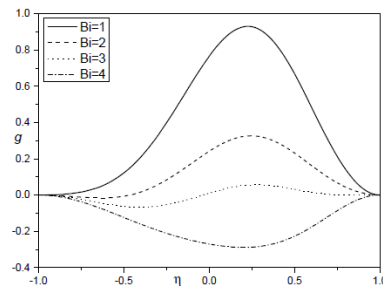


Fig 3. Effect of Bi on the cross-flow velocity

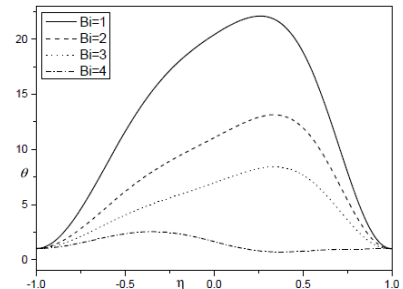


Fig 4. Effect of Bi on the temperature

The nature of $f(\eta)$, $g(\eta)$ and $\theta(\eta)$ at different values of $Ha, m, Da, \alpha, \beta_1$ and β_2 are presented through Figs. 5 to 22 by taking Pr, Re, Gr, ϵ at 0.71, 2, 2, 0.1 respectively.

The impact of m on $f(\eta)$, $g(\eta)$, and $\theta(\eta)$ can be identified in Fig. 5-7 at $\alpha=\pi/3, Br=1, Da=0.2, Ha=2, Bi=5, \beta_1=0.1$ and $\beta_2=0.1$. It is noticed from Fig. 5-7 that the flow velocity, cross-flow velocity, and temperature decrease with an increase in m . This is because of an inclined magnetic field. Here magnetic field applied with an angle $\alpha=\pi/3$ and Hall effect produces in the direction of inclined plates, hence it cannot act as a drag on velocities. As enlighten above the Hall current produces additional charge and which makes to decrease in temperature of the fluid.

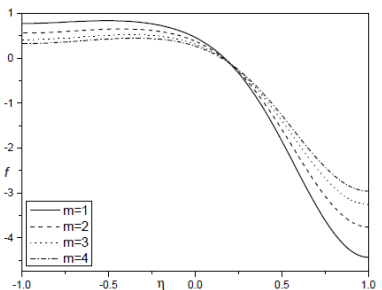


Fig 5. Effect of m on the flow velocity

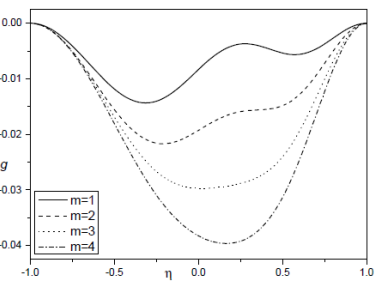


Fig 6. Effect of m on the cross-flow velocity

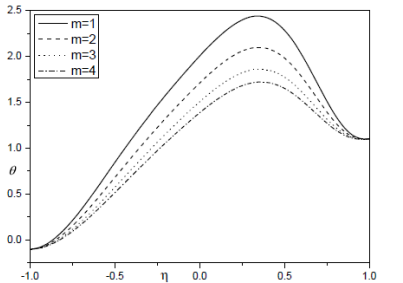


Fig 7. Effect of m on the temperature

Figure 8-10 presents the nature of f, g and θ under the influence of Darcy number (Da) when $\alpha=\pi/3, Br=2, m=4, Ha=2, Bi=5, \beta_1=0.1$ and $\beta_2=0.1$. It is noted from these figures that the flow velocities and temperature of the fluid increases with an increase in Da .

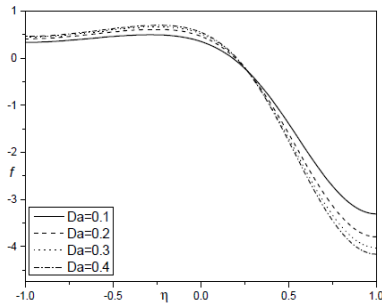


Fig 8. Effect of Da on the flow velocity

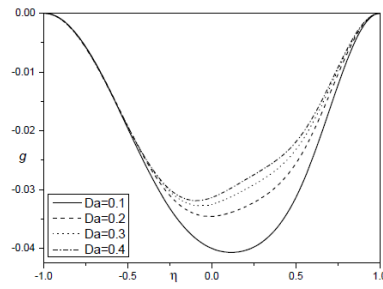


Fig 9. Effect of Da on the cross flow velocity

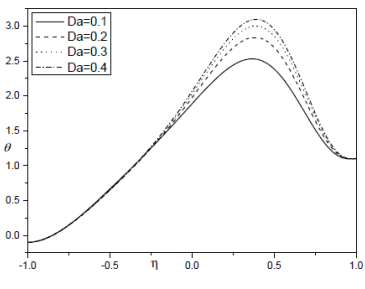


Fig 10. Effect of Da on temperature

Figure 11-13 presents the nature of f , g and θ under the influence of Ha when $\alpha=\pi/3$, $Br=1$, $Da=0.2$, $m=1$, $Bi=5$, $\beta_1=0.1$ and $\beta_2=0.1$. It is found in Fig. 11 that as Ha enhances the x -velocity increases. Here the magnetic field has an inclination angle $\alpha > 0$ applied to an inclined channel generates the drag force. It is identified from Fig. 12 that the cross-flow velocity increases as Ha increases. It can be shown from Figs. 13 that the temperature increases with the enhancement of the magnetic parameter. This is due to the reality that the magnetic field generates resistive power, which leads to an increase in the temperature.

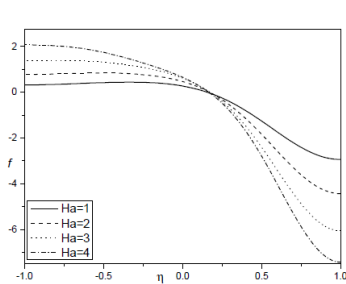


Fig 11. Influence of Ha on the flow velocity

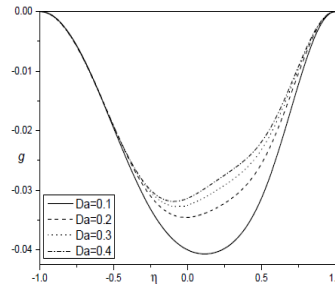


Fig 12. Influence of Ha on cross-flow velocity

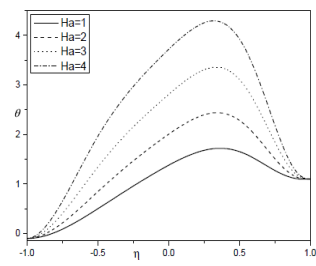


Fig 13. Influence of Ha on temperature

The impact of inclination angle α on f , g , and θ can be denoted in Fig. 14-16 by fixing the other parameters at $Da=0.2$, $Br=2$, $m=2$, $Ha=2$, $Bi=5$, $\beta_1=0.1$ and $\beta_2=0.1$. It is noticed from Fig. 14 and 16 that the flow velocity and dimensionless temperature increases as α enhances. It is observed in figure 15 that the cross-flow velocity decreases as α increases. This is due to the reality that as an inclination angle of applied magnetic field changes (angle of inclination increases) leads to the reduction in drag force will increase on the net flow.

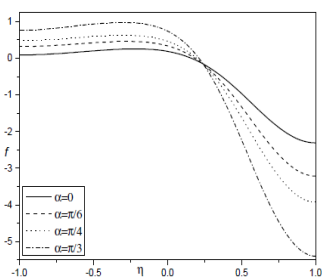


Fig 14. Effect of α on the flow velocity

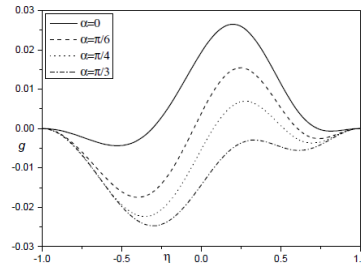


Fig 15. Effect of α on cross velocity

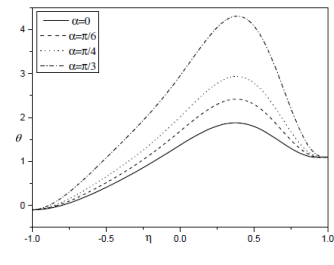


Fig 16. Effect of α on temperature

The effect of β_1 on f , g and θ can be noted in Fig. 17-19 by fixing the other parameters at $\alpha=\pi/3$, $Br=1$, $Da=0.2$, $m=2$, $Bi=5$, $Ha=2$ and $\beta_2=0.1$. It is observed from Fig. 19 that the temperature of the fluid increases as an

β_1 increases. It is noticed from Fig. 17 that the flow velocity decreases as β_1 increases. It is seen from Fig. 18 that the cross-flow velocity increases as β_1 increases.

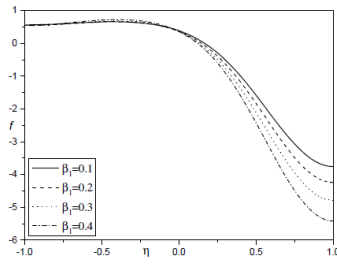


Fig 17. Effect of β_1 on $f(\eta)$

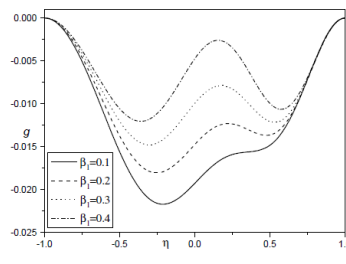


Fig 18. Effect of β_1 on $g(\eta)$

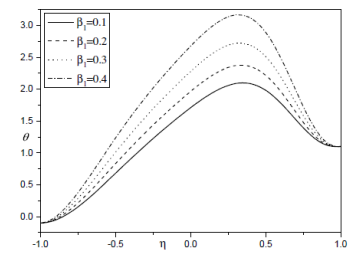


Fig 19. Effect of β_1 on $\theta(\eta)$

The effect of β_2 on f , g and θ can be noted in Fig. 20-22 by fixing the other parameters at $\alpha=\pi/3$, $Br=1$, $Da=0.2$, $m=2$, $Bi=5$, $Ha=2$ and $\beta_1=0.1$. It is noticed from Fig. 20 that the flow velocity decreases as β_2 increases. It is observed from Fig. 22 that the temperature of the fluid decreases as β_2 increases. It is noticed from Fig. 21 that the cross-flow velocity decreases as β_2 increases.

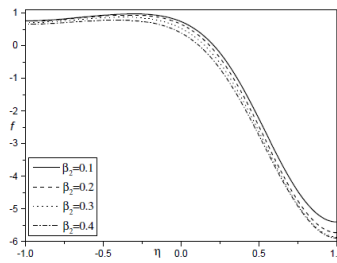


Fig 20. Effect of β_2 on $f(\eta)$

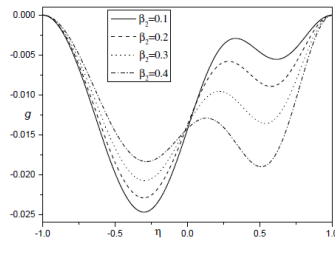


Fig 21. Effect of β_2 on $g(\eta)$

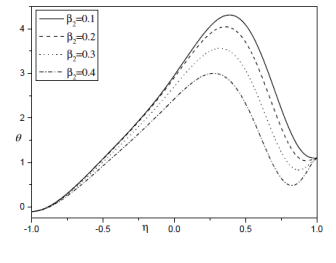


Fig 22. Effect of β_2 on $\theta(\eta)$

IV. CONCLUSION

This article investigated the steady inclined magnetohydrodynamic natural convective boundary condition in a porous channel in the occurrence of Hall and Darcy effects. Spectral Quasilinearization Method is used to solve the dimensionless governing equations. From this present study the main findings are listed as:

- The flow velocity profiles and temperature of the fluid decreases with the increase of Hall parameter (m), Biot number (Bi) and slip parameter (β_2).
- Flow velocity and temperature of the fluid are increases and cross-flow velocity profiles diminish with an enhance in inclination angle (α).
- The flow velocity profiles and temperature of the fluid increases with the increase of Hartmann number (Ha) and Darcy number (Da).
- Flow velocity decreases and cross-flow velocity and temperature are increases with the increase of slip parameter (β_1).

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